

## **GNSS signals as a tool for detection of the influence of solar radiation of terrestrial ionosphere**

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Solar radiation has the most significant influence on photo-ionization in the Earth's ionosphere, and different parts of its spectrum in the UV and X domains have a dominant role in producing free electrons at different ionospheric heights. In calm conditions, the dominant role in photo-ionization processes in the D-region has the hydrogen Ly $\alpha$  line, while in the higher layers, the ionosphere is primarily ionized by photons from the X spectrum.

Through ionization processes, solar radiation significantly affects changes in the electron density in the ionosphere and, consequently, the propagation of electromagnetic (EM) waves in this atmospheric layer. One of the most significant EM waves propagating in the ionosphere is Global Navigation Satellite Systems (GNSS) signals with various applications, including navigation and positioning applications. Since 1987 International Association of Geodesy (IAG), as one of the several associations of the International Union of Geodesy and Geophysics (IUGG), has been responsible for establishing and maintaining European Terrestrial Reference System (ETRS). To establish and maintain ETRS, IAG has formed a special commission entitled EUREF (European Reference Frame), responsible for creating and maintaining the EUREF Permanent Network (EPN) as a science-driven network of continuously operating GNSS reference stations. More than 100 European agencies and universities are voluntarily involved in the EPN, with 368 stations all over Europe. In general, all GNSS stations continuously receive the data from all currently available GNSS: American NavStar, Russian GLONASS, Chinese BeiDou, and European Galileo systems. The received data are stored in EPN Data Centres (DC), and all data are publicly available in several RINEX formats (Receiver Independent Exchange Format). Besides data centers, EUREF organized EPN Analysis Centers (AC) to provide products such as the EPN GNSS station positions and velocities, EPN zenith tropospheric path delay estimates, and ETRS89 satellite orbit and correction streams.

In the precise positioning process, the crucial element is modeling the ionosphere, which gives significance for practical applications to studies that analyze the influence of solar radiation and its spectrum on photo-ionization processes. Namely, the ionosphere delays EM waves that pass to it and induces a

delay of propagation related to the electronic density (Total Electron Content - TEC) and the wave frequency. This delay is directly proportional to the TEC and inversely proportional to the square of the carrier frequency. Those effects can produce an error in points coordinates from 3 to 15 m. Besides this effect, there are also effects caused by irregularities in the ionosphere – scintillation effects – which can cause a large number of cycle slips in GNSS signals. The magnitude of mentioned errors is the main reason for many studies and papers dedicated to modeling ionospheric effects on GNSS and using GNSS signals to determine global, regional, and local ionospheric models.

It is usual in GNSS applications to split the determination of ionosphere into deterministic and stochastic parts. The deterministic part is dedicated to creating models of the ionosphere, and the stochastic part is dedicated to the determination of short-term TEC variation throughout the estimation of Stochastic Ionosphere Parameters (SIP). This paper is dedicated to determining the deterministic part by modeling local, regional, and station-specific ionosphere models. All mentioned models are created based on available EPN data by Bernese GNSS Software Version 5.2 and mentioned models are represented by Taylor series expansion in the case of local models and spherical harmonic expansion for the regional and station-specific models. All models can be produced for the desired period and, in such a way, create relevant scientific data for studying the effect in the ionosphere induced by changes in solar radiation.