

INFRARED EMISSION OF THE AGN DUSTY TORUS: RADIATIVE TRANSFER MODELING WITH SKIRT

M. Stalevski^{1,2}, J. Fritz², M. Baes² and L. Č. Popović¹

¹*Astronomical Observatory, Volgina 7, 11060 Belgrade, Serbia*

²*Sterrenkundig Observatorium, Universiteit Gent, Krijgslaan 281-S9,
Gent, 9000, Belgium*

E-mail: mstalevski@aob.rs

We investigated the infrared emission of a toroidal structure of gas and dust (the “dusty torus”) surrounding the central regions of the AGNs. We introduced a new model of the dusty torus as a clumpy two-phase medium, with high-density clumps embedded in a low-density interclump dust. To obtain spectral energy distributions and images of the torus at different wavelengths, we employed the 3D Monte Carlo radiative transfer code SKIRT. We calculated a grid of models for different parameters, analyzed the properties of infrared emission and compared them to the properties of the corresponding sets of clumps-only models and models with a smooth dust distribution. We found that the most striking feature of the two-phase model is that it might offer a natural solution to the common issue reported in a number of papers — the observed excess of the near-infrared emission. Namely, a torus model with the dust distributed in a two-phase medium has a more pronounced (‘hotter’) emission in the 2 – 6 μm range while displaying, at the same time, an attenuated 10 μm silicate feature.