



# The Connection Between Broad Emission Lines Properties and Stellar Velocity Dispersion in Sample of AGNs Type 1

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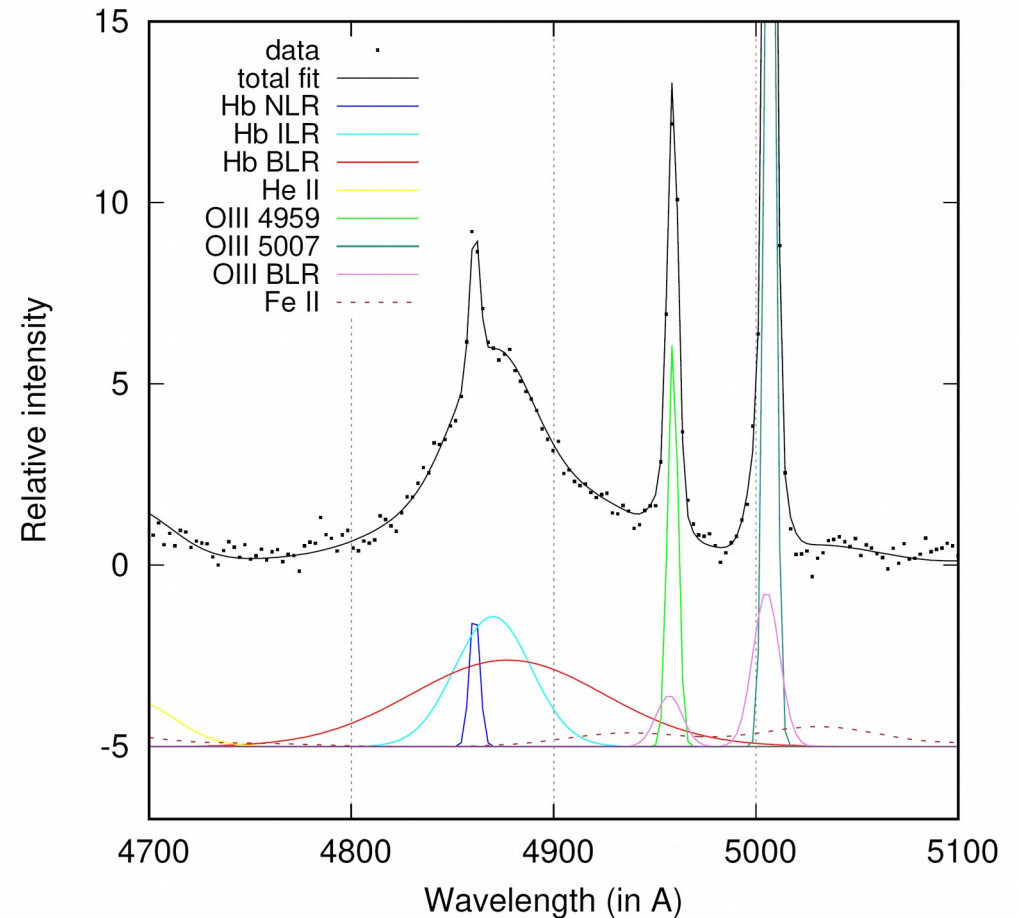
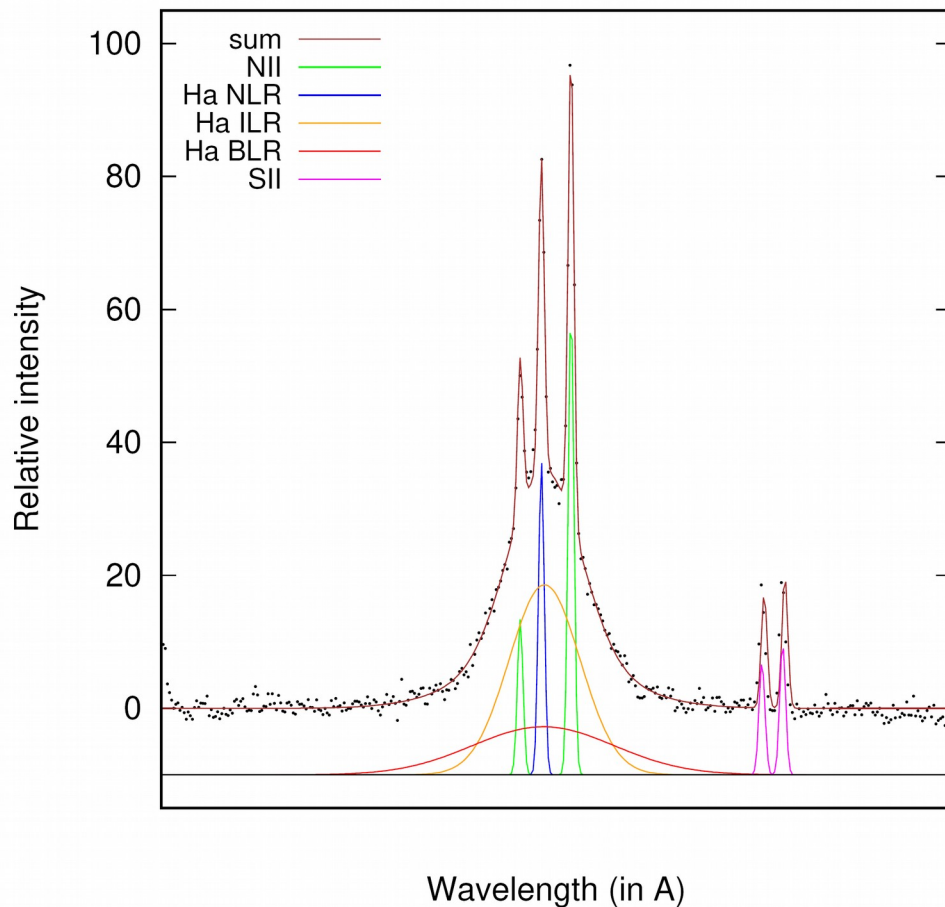
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# Spectroscopic estimation of SMBH mass

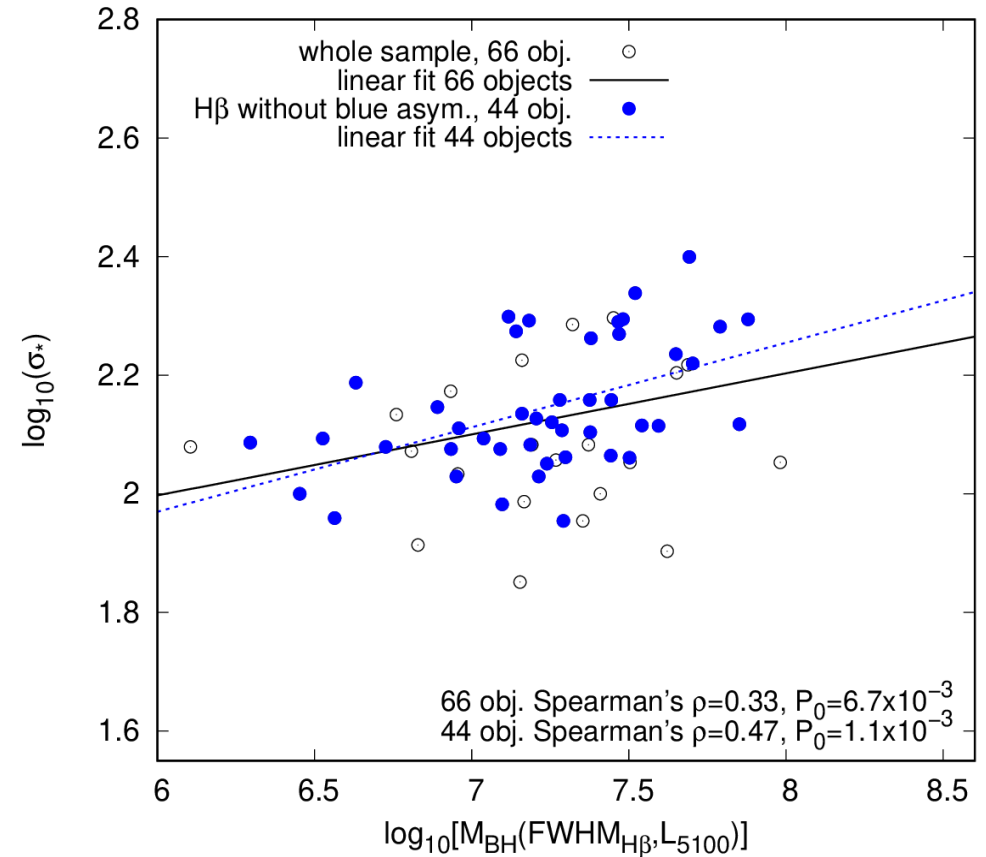
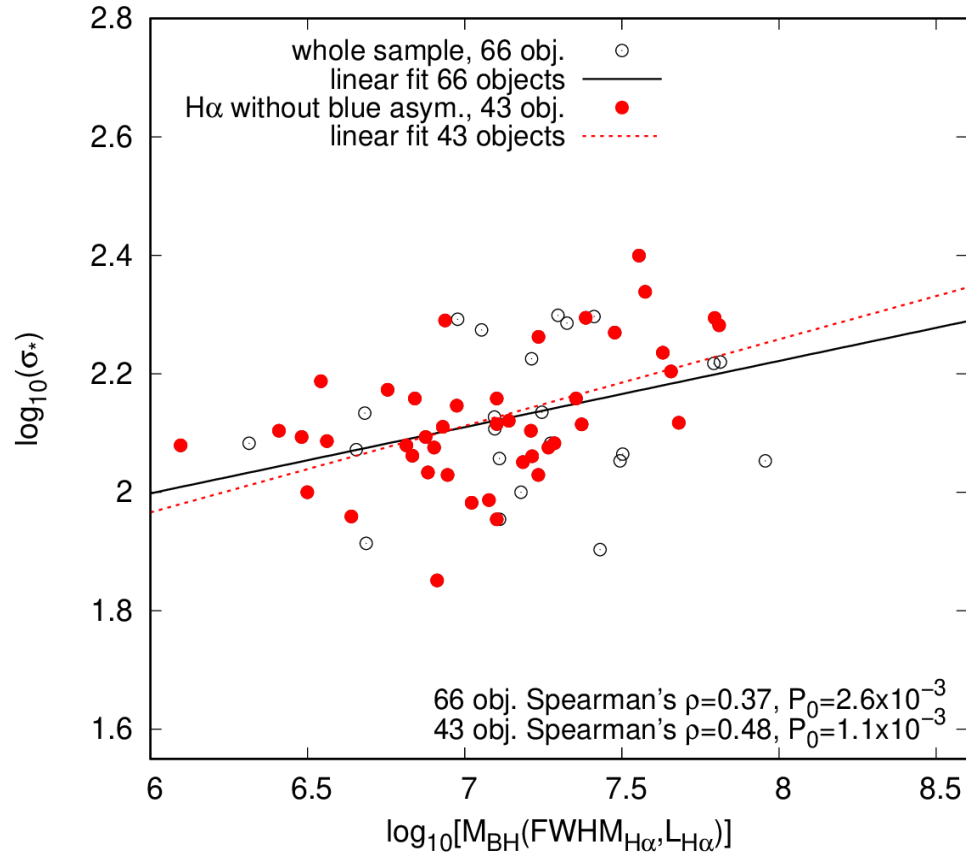
- Several methods for estimating the super-massive black hole (SMBH) mass  $M_{\text{BH}}$  use spectral properties of AGN emission lines.
- One of the most common method is single-epoch virial method, that uses continuum luminosity at 5100nm + full width at half maximum (FWHM) of the broad  $\text{H}\beta$  emission line (Peterson, 2014)
- There is an assumption that the broad  $\text{H}\alpha$  emission line also could be used for estimation of  $M_{\text{BH}}$  (Green&Ho, 2005)
- Here we compare virial masses obtained from  $\text{H}\alpha$  and  $\text{H}\beta$  lines, with stellar velocity dispersion ( $\sigma^*$ )

# Sample and Analysis

- We used the sample of 66 AGNs Type 1, with available  $\sigma^*$ . Spectra were taken from the SDSS, analyzed in Shen+2015 and in Harris+2012
- Spectra were corrected for the Galactic extinction, cosmological redshift, host galaxy contribution, continuum emission and then fitted using the multi-Gaussian model of optical emission (Kovačević+2010)

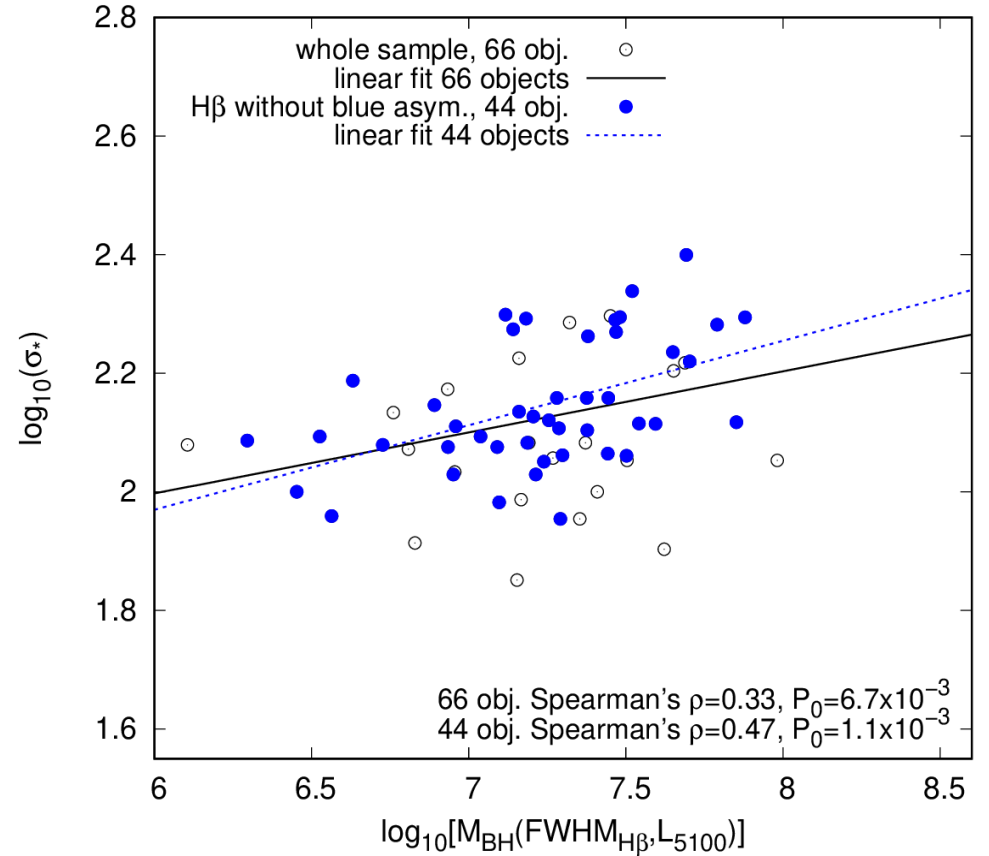
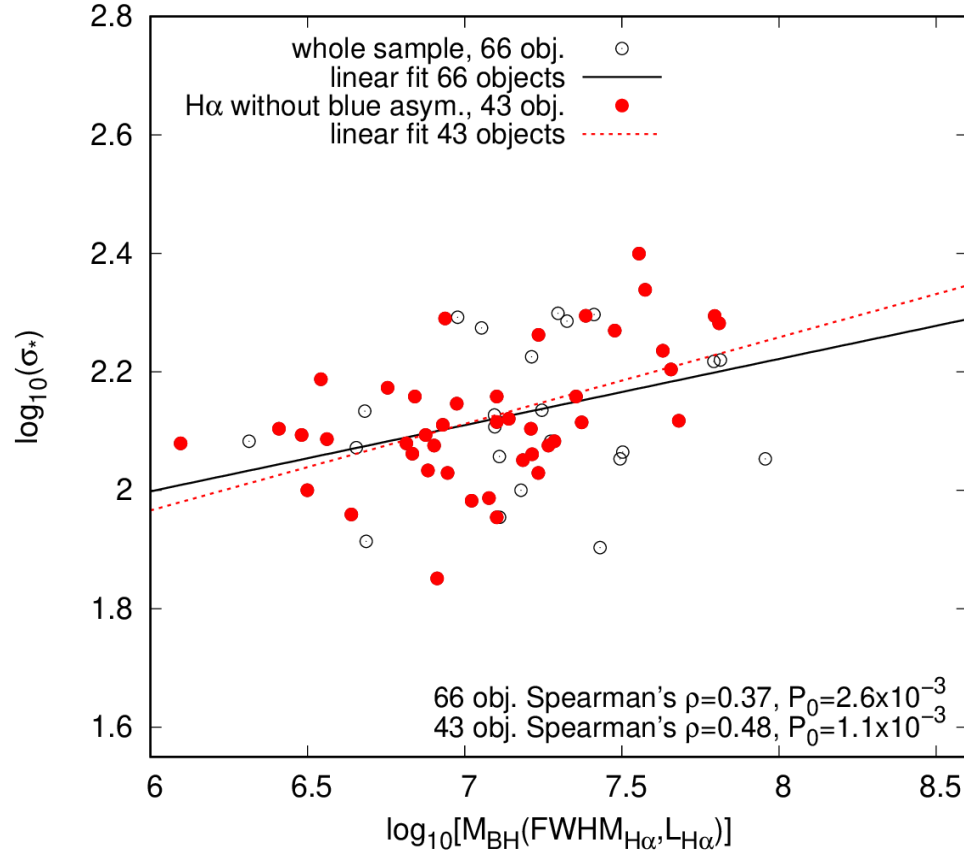


# Results



- Here we present comparisons of  $\sigma^*$  with virial  $M_{\text{BH}}$  obtained using H $\alpha$  (left) and H $\beta$  (right)
- We supposed that blue asymmetry in BEL indicates non gravitational effects, and that removing these object from the sample would improve these correlations. In order to check that, we made subsamples of objects having only red asymmetry in emission lines (full circles).

# Conclusions



- We confirmed that parameters of the broad  $H\alpha$  emission line are reliable for  $M_{BH}$  determination, as those from  $H\beta$  are.
- In both cases, the correlation of  $\sigma^*$  and SMBH mass obtained by the virial method, are significantly improved when the objects with blue asymmetry in emission line, were removed from the sample (for  $H\alpha$  from  $\rho=0.37$  to  $\rho=0.48$ ,  $P_0=10^{-3}$ , and for the  $H\beta$  from  $\rho=0.33$   $\rho=0.47$  and  $P_0=10^{-3}$ ).

*Thank you*